

The new outburst of the Be star ω Ori

© Ernst Pollmann

Emil Nolde Straße 12 - 51375 Leverkusen



4. Mai 2006

1) The Be star ω Ori entered in strong outburst. It apparently started in september. The $H\alpha$ -line is currently at 2 times the continuum in emission.

The FeII lines are also in emission and visible all around the spectrum. It is thus a very interesting spectrum to take and a very interesting star to follow in the coming weeks and months.

One can read Neiner et al. 2002 (A&A 388, p. 899) to know everything about the star, its pulsations and the long-term evolution of its emission.

One can also read Neiner et al. 2003 (A&A 409, p. 275) to know everything about the rotation and magnetism of this star.

C.Neiner

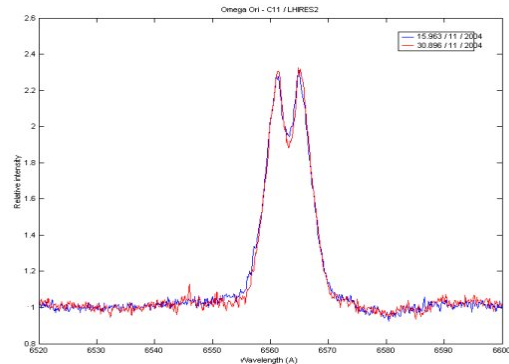
2) Graphical form of ω Ori spectra for the date 15/11/2004 and 30/11/2004. Low evolution is noticed at the resolution of $R=16\ 000$.

The measured equivalent width for the 30/11 observation is $10.0 \pm 0.2 \text{ \AA}$, very similar to the Ernst Pollmann evaluation (10.3 \AA).

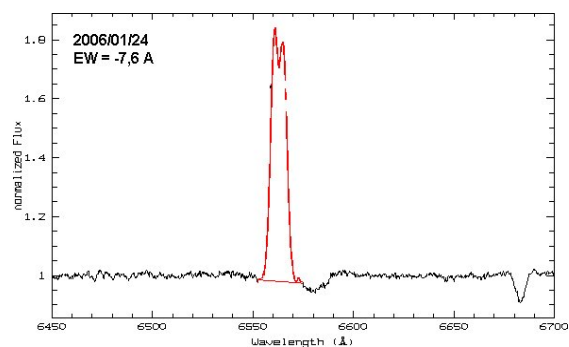
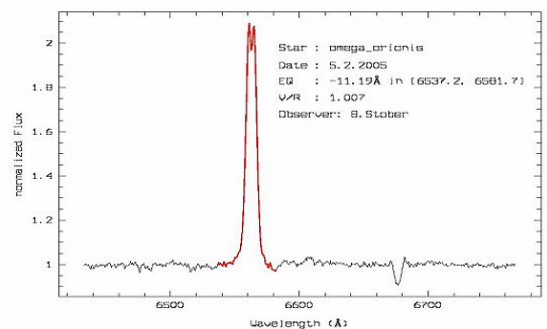
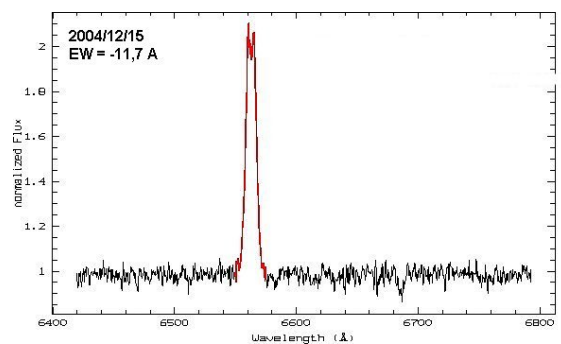
Ch.Buil

3) $H\alpha$ -spectra by Ch. Buil and the VdS-spectroscopy group (Germany)

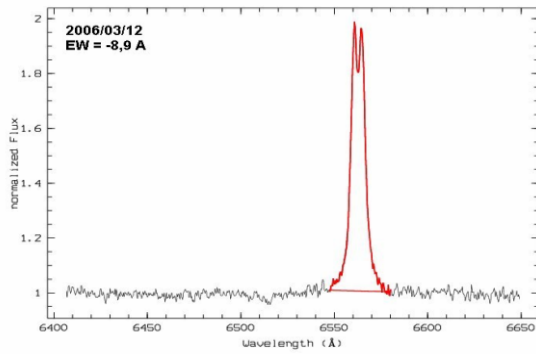
a) Spectra by Ch. Buil



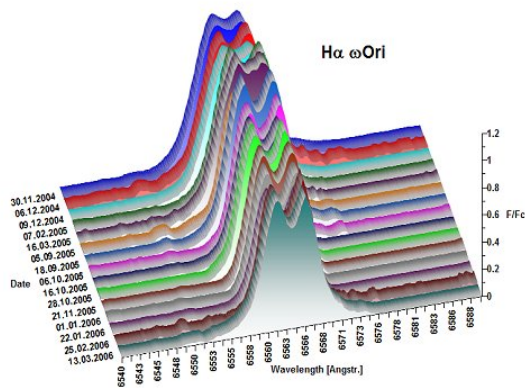
b) Spectra by B. Stober



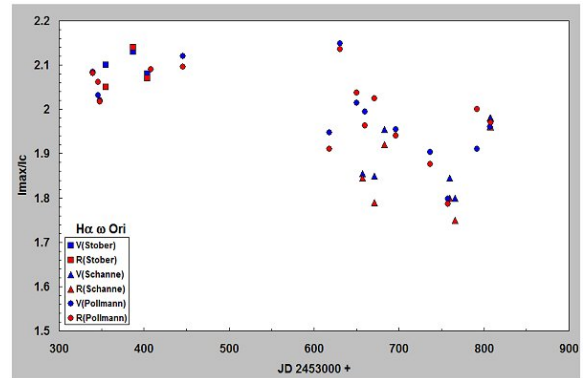
c) Spectra by L. Schanne



d) Spectra by E. Pollmann



5) Measurements of the intensity maxima of the V- and R-component of the H α -peak



4) Measurements of the H α -Äquivalentwidth

E. Pollmann: 20 cm Schmidt-Cassegrain Telescope; Dispersion: 0.25 Å/Pixel
B. Stober: 30 cm Newton Telescope; Dispersion 0.6 Å/Pixel
L. Schanne: 15cm Telescope; Dispersion 0.6 Å/Pixel

